

High-Speed Atmospheric Correction Algorithms for Spectral Image Processing*

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** NASA SBIR Phase I, D. Mandl (NASA GSFC TPOC)*





Topics

- **SSI Overview**
- **VNIR-SWIR Atmospheric Correction**
- **Thermal IR Atmospheric Correction**
- **MODTRAN**
- **MCSce (3D Scene Simulation)**

Spectral Sciences, Inc. Overview



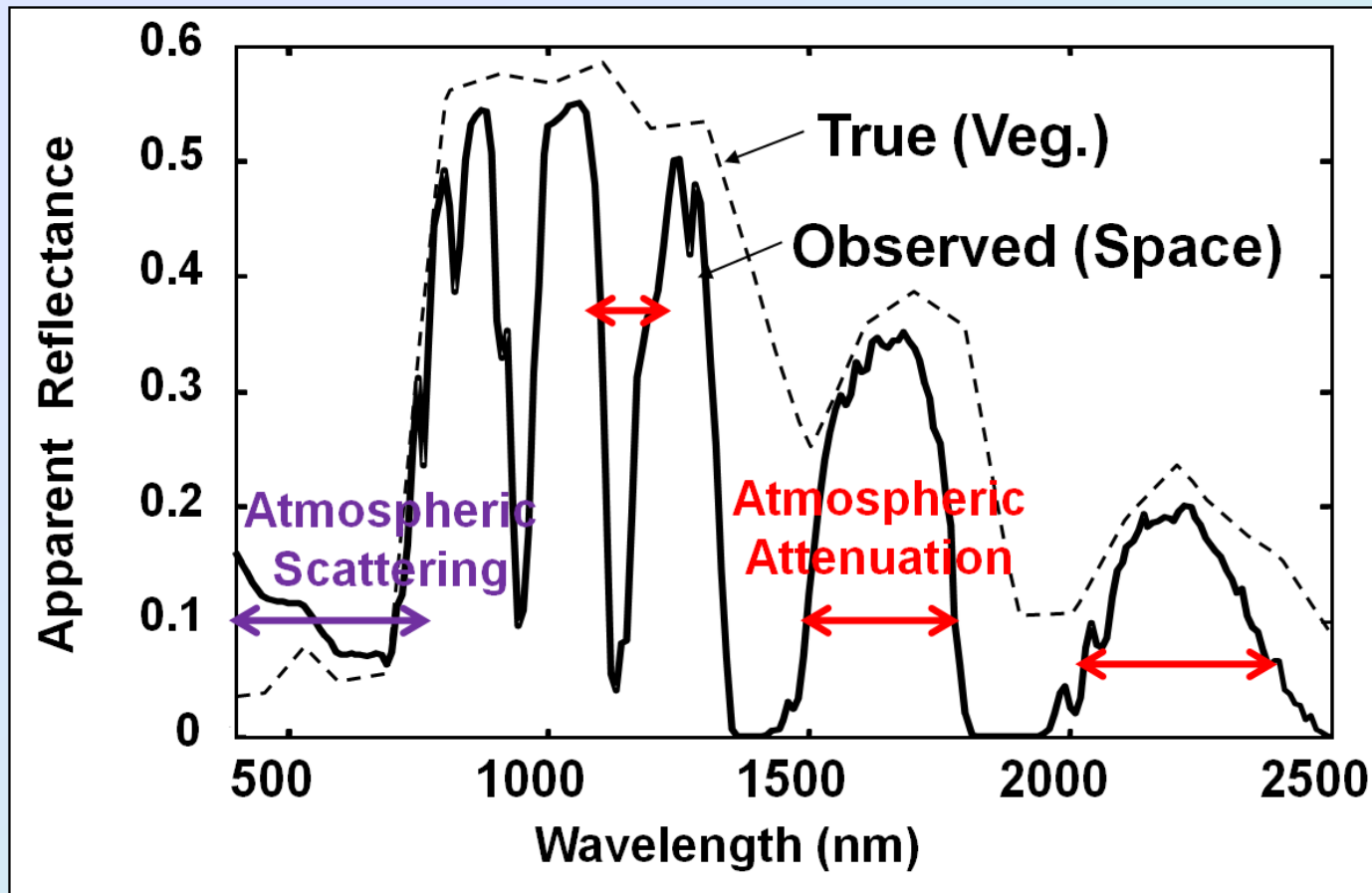
- ***Incorporated: January, 1981***
 - ***Near to Burlington Mall***
- ***Annual Sales: \$10M***
- ***Full-Time Staff of 42***
 - ***33 Ph.D. and 2 M. S.***
- ***Major Business Areas***
 - ***Remote Sensing, Physics Modeling, Basic Research & Software Development***
 - ***Field Experiment Planning and Data Analysis***
 - ***Trace Gas/Industrial Sensor Development***

We apply basic scientific research to challenging problems in the National Defense and Commercial sectors.



Atmospheric Correction Basics

Scattered Sunlight off Earth Surface



**Observed
Signal**
↓
**Atmospheric
Correction
Algorithm**
↓
**True
Spectral
Reflectance**



FLAASH Overview

- Development
 - SSI lead with primary support from AFRL, additional support from NGA, NASA, SSI
 - ENVI commercial product with ITT-VIS
 - FLAASH-C developed for NGA parallel processing system
- Science/Features
 - MODTRAN radiative transfer; pixel-by-pixel water retrieval; scene visibility retrieval; adjacency effect and spectral smile compensation; spectral polishing; wavelength self-calibration
- Operating Modes
 - Interactive (IDL) or batch (FLAASH-C)
 - High-speed MODTRAN LUT option (ongoing development)
- Demonstrated Sensor Support
 - AISA-ES, ALI, ARTEMIS, ASAS, ASTER, AVHRR, AVIRIS, CASI, Compass, GeoEye-1, HYDICE, HyMap, Hyperion, IKONOS, Landsat, LASH, MaRS, MASTER, MODIS, MTI, Probe-1, QuickBird, RapidEye, SPOT, TRWIS, WorldView-2



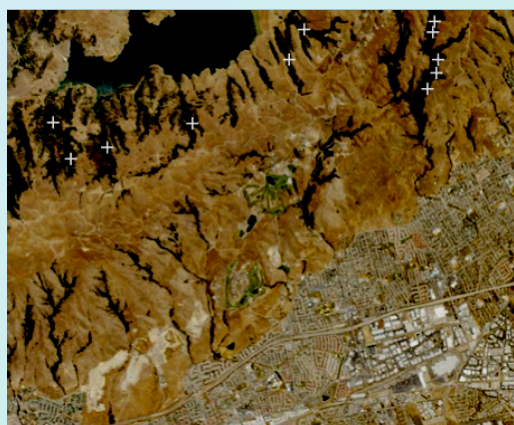
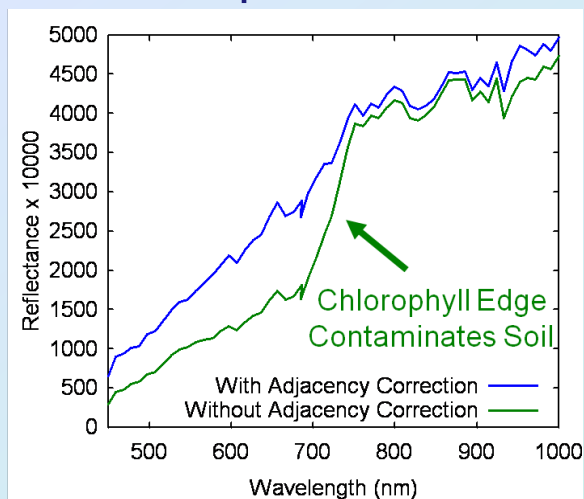
FLAASH Science

RT Equation

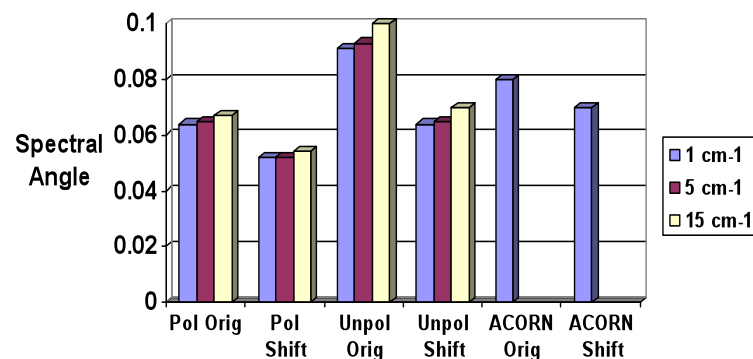
$$L^* = \frac{a \rho}{1 - \rho_e S} + \frac{b \rho_e}{1 - \rho_e S} + L_a^*$$

MODTRAN-derived

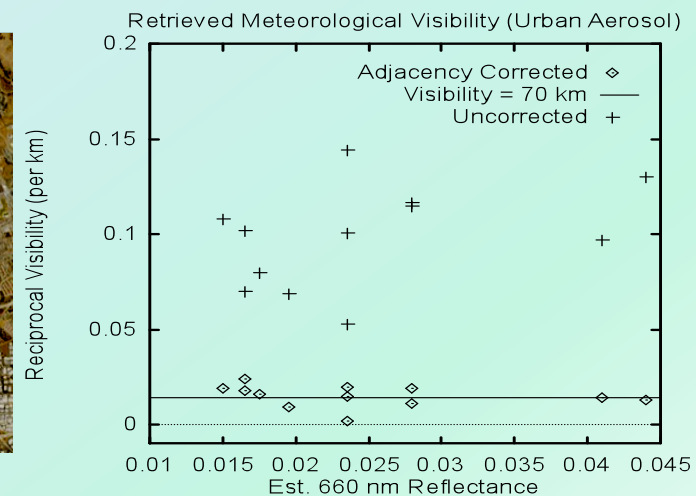
Adjacency Compensation



Validation with Ground Truth



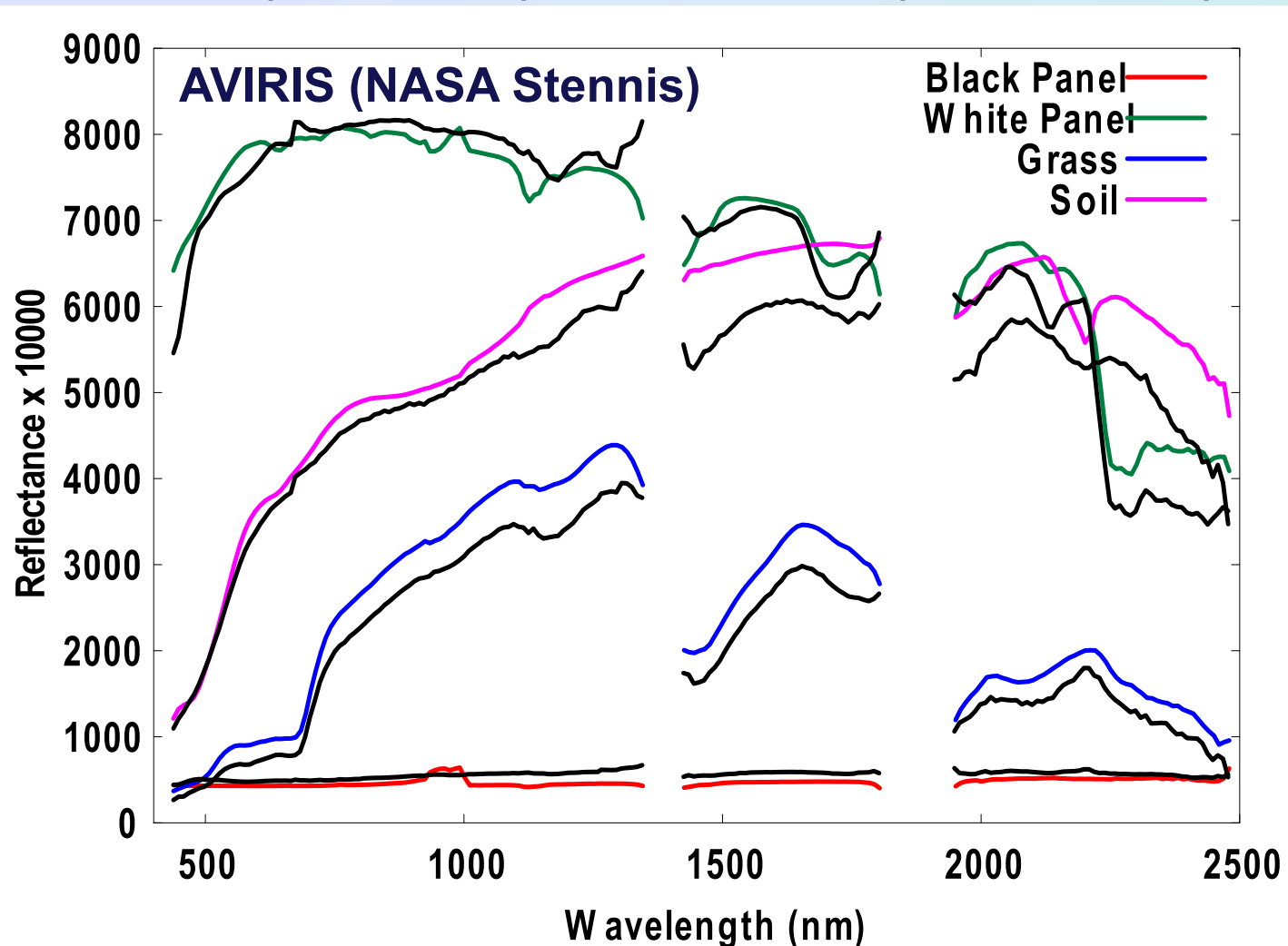
Visibility Retrieval





FLAASH Comparison to “Truth”

FLAASH (black lines) ASD Spectra (colored lines)





Current Phase I SBIR Program

- Problem
 - High duty cycle of upcoming missions (e.g., HypsIRI) requires high accuracy, fully automated, low latency, near real-time atmospheric correction (AC) processing
 - NASA's current AC algorithms require upgrades in science and coding
- Solution
 - Transition NASA's current and future AC processing to a fast version of the C++ language FLAASH-C code
- SBIR Program Objectives
 - Port FLAASH to the Elastic Cloud or IAAS
 - Develop a look-up table (LUT) for near-real-time FLAASH processing
 - Phase II: implement on prototype flight hardware
- Team: Spectral Sciences, Inc. (S. Adler-Golden) and Vightel Corp. (P. Cappelaere)



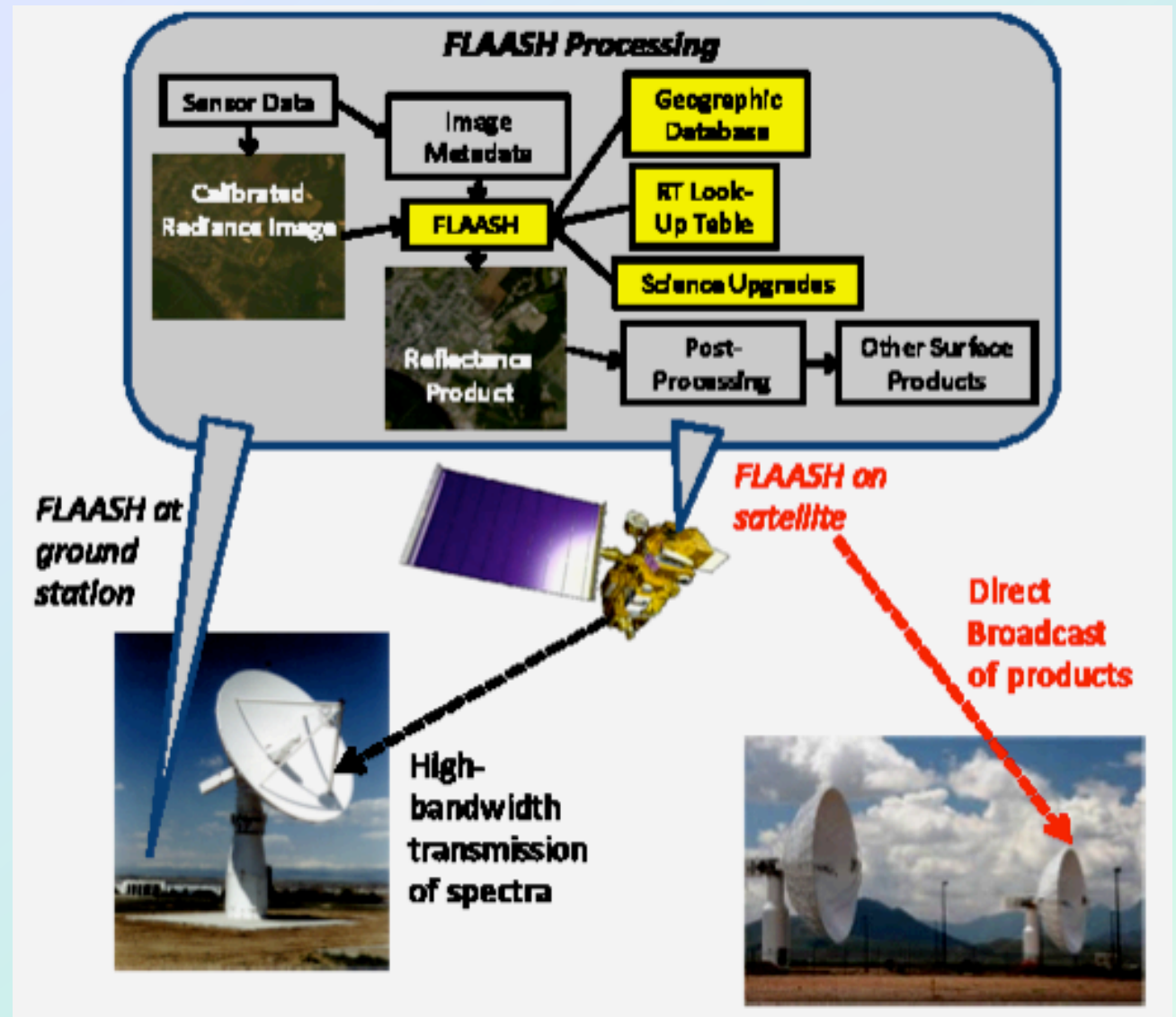
Processing Concept

- **Ground-based system:**

- Will support Hyperion, Landsat, ALI, MODIS, ASTER

- **On-board system:**

- For direct broadcast of products from HypIRI, LDCM in near real-time



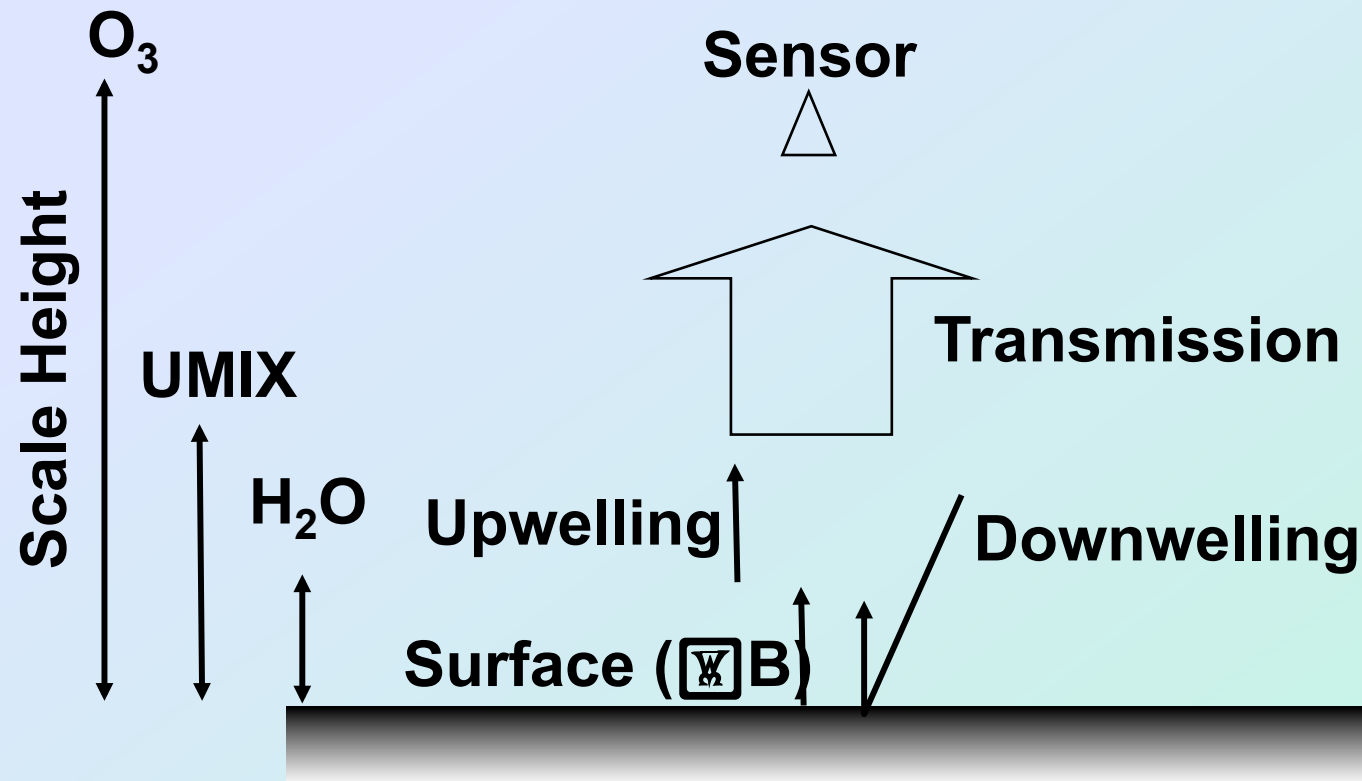


FLAASH Speedup via Look-up Table (LUT)

- Large LUT replaces the custom MODTRAN calculations
 - Eliminates Fortran module and results in a *~2x Speed Up*
 - FLAASH-C timing with LUT - *~ 10 sec* for Hyperion data strip using a single 2.7 GHz micro-processor
- In feasibility demo in a 2000 NASA Ph 1 program
 - “Giant” LUTs (GLUTs) built for AVIRIS (20 km alt) and HYDICE (3 km alt) data
- SSI recently delivered FLAASH-C with the 20 km GLUT to Goddard for testing
 - Installing on Elastic Cloud system
- In Phase I new LUTs are being build for space applications
 - TOA (Top Of the Atmosphere) nadir and off-nadir
- Storage: original GLUT *~ 615 Mb*; new LUT database will require compression — e.g., *~10x* (slightly lossy) with SVD



Infrared Atmospheric Correction



Goal: Remove atmospheric effects and retrieve surface emissivity ϵ and temperature T_s .



IR Atmospheric Correction Algorithms

- **Underlying problem** – more unknowns than knowns, thus one needs to introduce a physically meaningful constraint
- **Spectral Smoothness (C. Borel)**: Vary atmosphere and T_{surf} to retrieve spectrally smoothest emissivity ϵ
 - Limited to Hyperspectral data
- **Minimum T_{surf} Variance (SSI & NG)**: Vary atmosphere and ϵ until narrowest distribution of T_{surf} is obtained
 - Applicable to Hyper- and Multi-spectral data
- **Minimum T_{surf} Variance with Endmember Expansion (SSI & LM)**: Only need to correct endmembers
 - Very fast
- Fully automated ENVI/IDL prototypes for all three methods



MODTRAN5 Overview



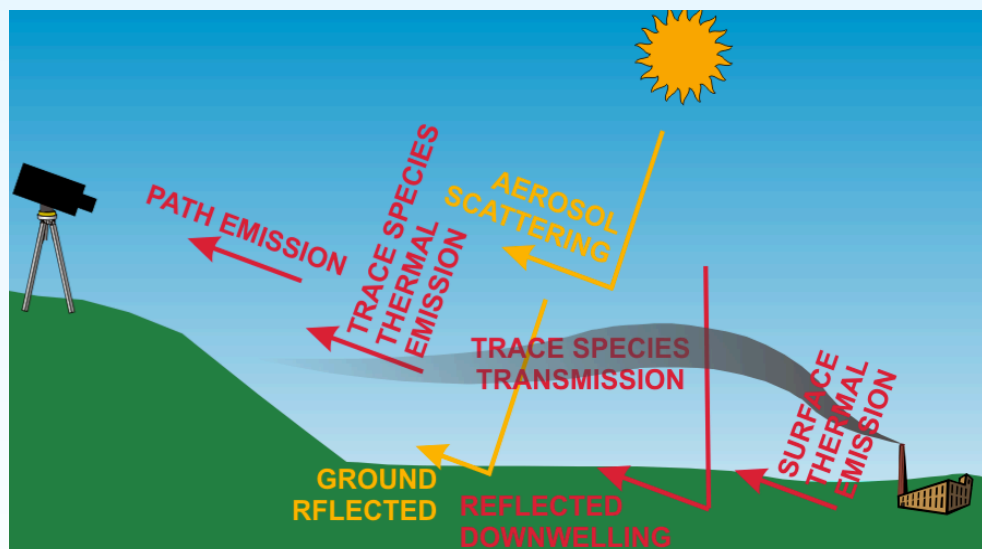
- Overview

- IR-Visible-UV Spectral and Channel Transmittances, Radiances, Fluxes, and ... from 0.1, 1.0, 5.0 or 15.0 cm^{-1} Band Models
- Stratified (1D) Molecular / Aerosol / Cloud Atmosphere
- 2-Stream and DISORT Solar and Thermal Scattering
- Spherical Refractive Geometry

- Applications

- Sensor Design
- Atmospheric Correction
- Measurement Planning
- Data Analysis
- Scene Simulation
- Algorithm Development
- **Local chemical plume
PLUG IN for MODTRAN5.3**

- **PNNL Data Base**

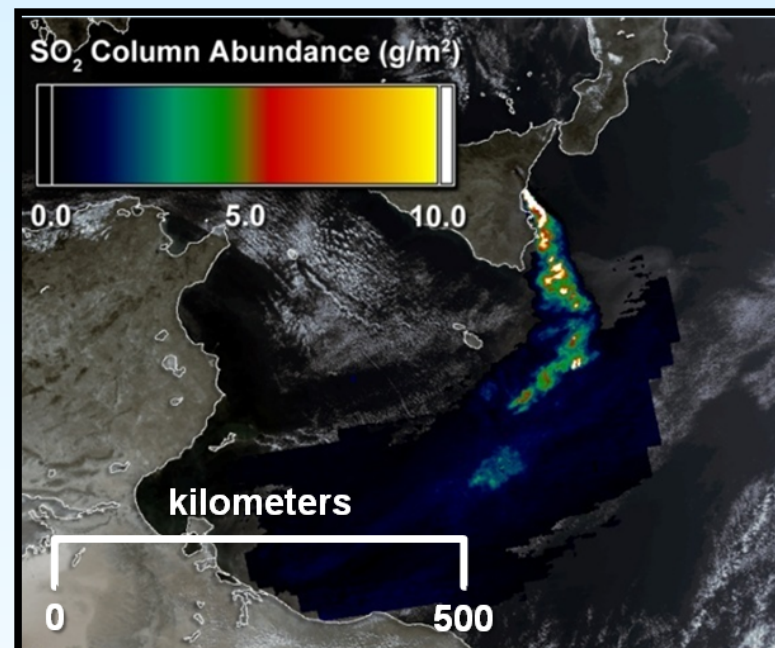




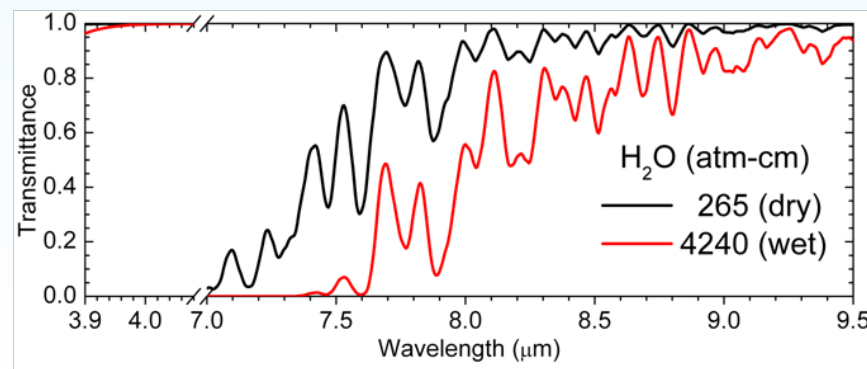
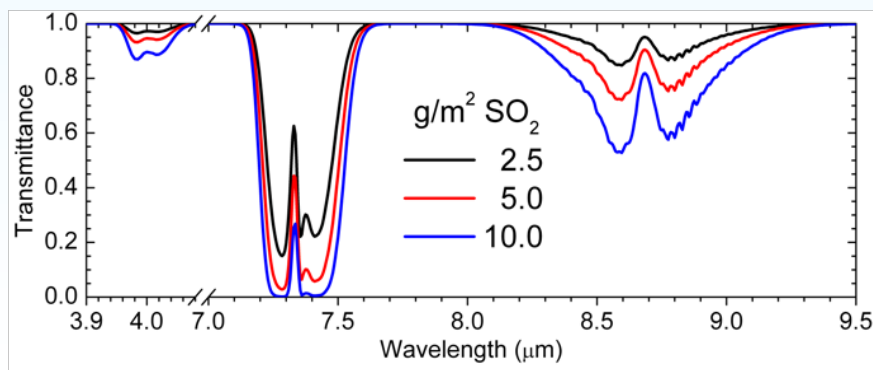
Remote Sensing of Volcanic Plumes from Space



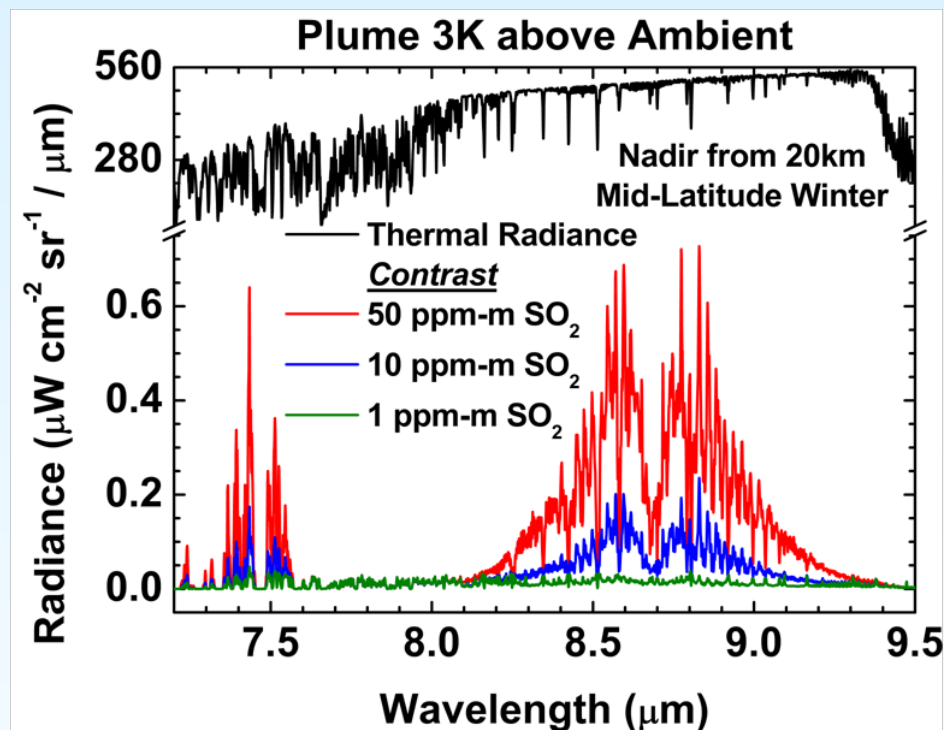
- The NASA Decadal Survey suggests that impending volcanic eruptions may be signaled by changes in gas emissions rates
- SO_2 column abundances $>1\text{g/m}^2$ (~ 350 ppm-m) have been derived from MODIS data
- MODTRAN local gas plume option provides a tool for modeling SO_2 contrast signatures
 - H_2O absorbance interferes with the stronger SO_2 absorption bands



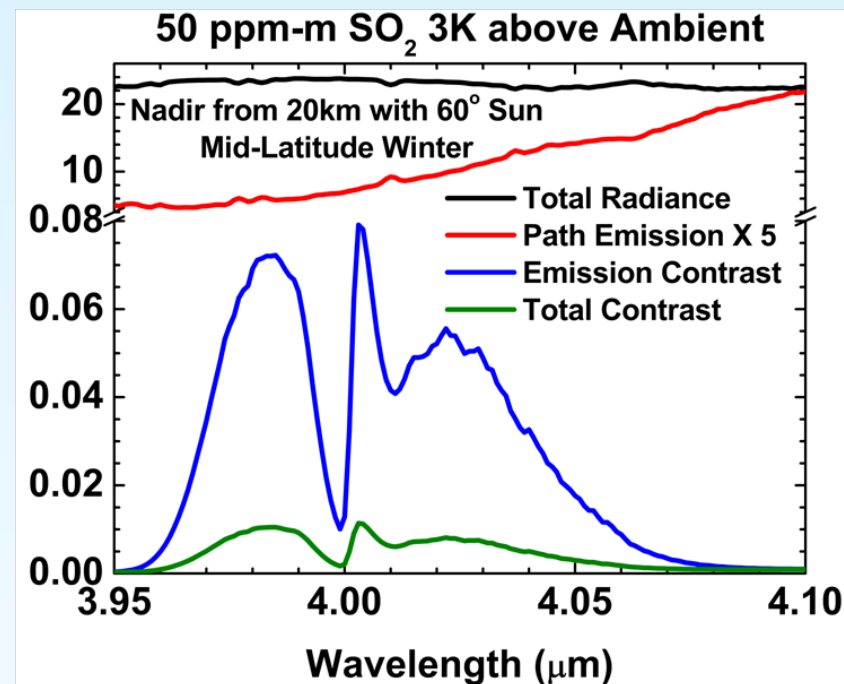
Klyuchevskoy Volcano
MODIS Retrievals of SO_2 column



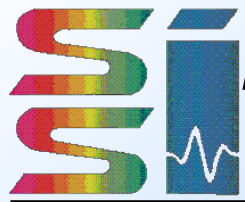
Simulated SO_2 Release



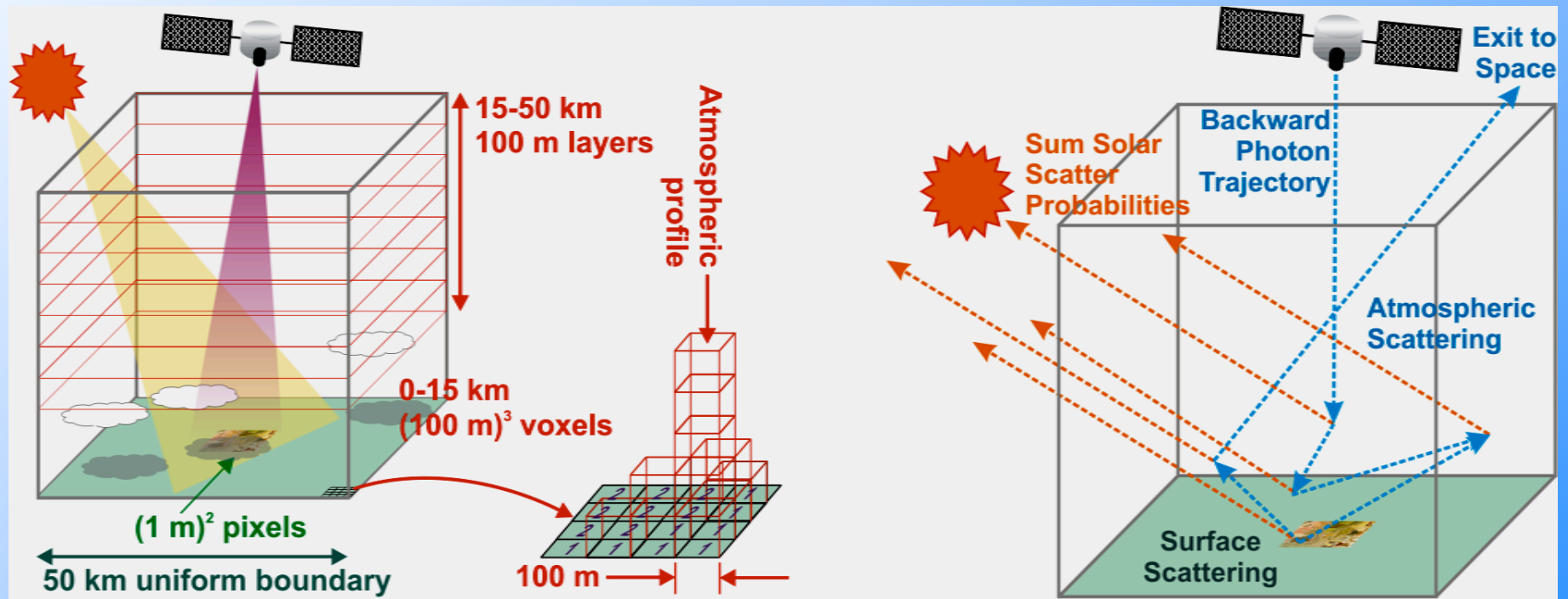
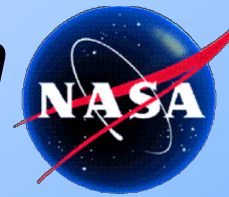
- Long-wave sensitivity study using MLOS option
- Contrast includes thermal emission and a small thermal scatter component



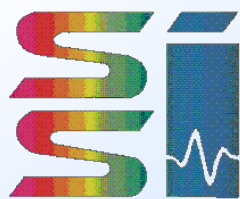
- Near 4 μm , scatter component is ~an order of magnitude larger than the thermal
- Ignoring the scatter contrast would produce a large error



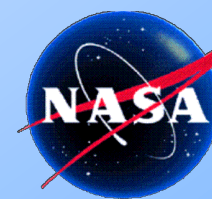
MCS Scene Reflective Domain Approach



- 3D atmospheres with 3D clouds and cloud fields
- 3D surfaces with BRDF's and object insertion
- 3D faceted object insertion
- **MODTRAN** optical properties, profiles, and spectroscopy



Simulation of Visibility

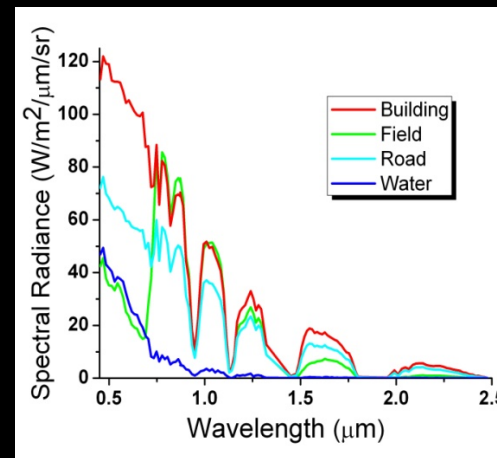


- *Davis CA scene*
 - Reflectance from HYMAP data
 - 121 spectral bands
 - 2.75 m GSD
- *Sensor altitude*
 - 50 km
 - nadir view
- *45 deg solar zenith angle*
- *Uniform atmosphere*
 - Mid-latitude summer
 - Rural aerosol
- *Reflectance outside of scene is assigned the in-band average*

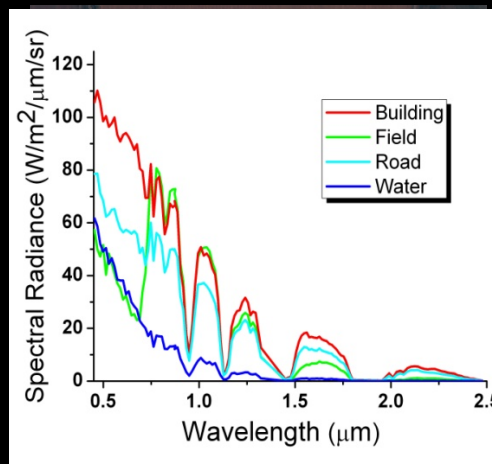
no atmosphere



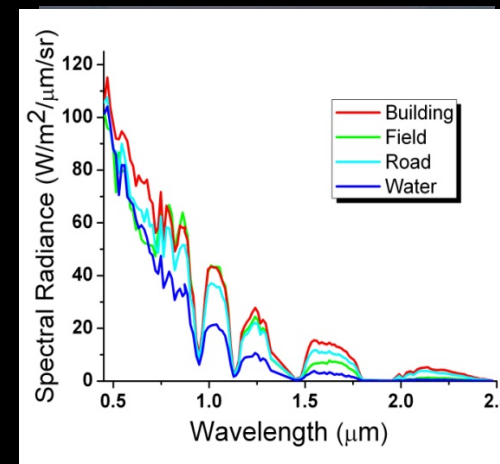
vis=100 km

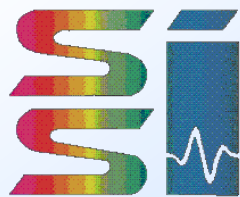


vis=23 km



vis=5 km



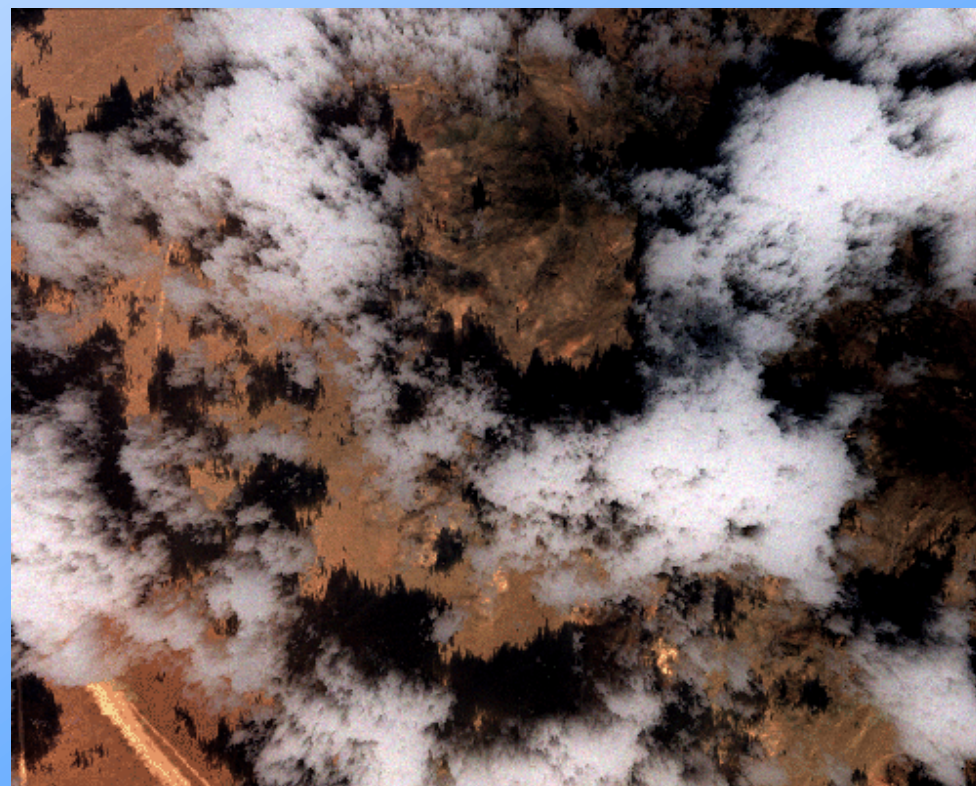


CSSM*

(Cloud Scene Simulation Model)



- *Realistic high-resolution cloud features, defined by larger-scale weather conditions in 4D*
- *Relies on stochastic field generation techniques and convection physics*
- *3D cloud voxels contain cloud properties which are used in MCScene*
 - *Water state*
 - *Water density*
 - *Precipitation density*



*Cianciolo and Raffensberger, "Atmospheric Scene Simulation Modeling and Visualization (AMV): Cloud Scene Simulation Model User's Guide," TASC Report TIM-07169-2,(1996).



Thank You

Any Questions?